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Section A

Light concentrators for the Sudbury Neutrino Observatory

G. Doucas[a], S. Gil[b,1], N.A. Jelley[a,*], L. McGarry[b], M.E. Moorhead[a], N.W. Tanner[a], C.E. Waltham[b]

[a] *Nuclear Physics Laboratory, Oxford University, Keble Road, Oxford, England OX1 3RH, UK*
[b] *Physics Department, University of British Columbia, Vancouver B.C., Canada V6T 1Z1*

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Abstract

There is an important and growing class of elementary particle detectors which are characterized by a large sensitive volume (thousands of tonnes), very low radioactive backgrounds, and rely on the emission of light for particle detection. Water Cherenkov detectors come into this category; they have a large mass of water as the sensitive medium. Particles are detected when they interact with the water and produce Cherenkov light, so detection efficiency relies on having a huge light sensitive area at the periphery of the detector. The most cost-effective way of achieving this is by placing light concentrators on large photomultiplier tubes (PMTs). This paper describes the work carried out on light concentrators for the PMTs in the Sudbury Neutrino Observatory, a 1000 tonne heavy water Cherenkov detector. We discuss the advantages of using light concentrators, summarize the optical theory of non-imaging light concentration, and describe in detail the development and manufacture of the concentrators themselves.

1. Introduction

The light concentrators we consider are devices which take light incident on an entrance aperture over a given angular range, and direct it by reflection from a curved surface onto a smaller exit aperture. There is thus a gain in collection efficiency with a corresponding reduction in angular coverage in accordance with phase space volume conservation. The Winston cones on the end of light guides are an example of light concentrators familiar to particle physics. The concentrators we describe in this paper are a development of the Winston cone for use on spherical PMTs.

1.1. Large water Cherenkov detectors

When elementary charged particles pass through a medium at a speed which is faster than the local speed of light, they radiate "Cherenkov light" [1]. These photons have a $1/\lambda^2$ wavelength distribution, but those with $\lambda < 300$ nm are usually absorbed in detector components, so the useful light is predominantly in the blue and near UV. In water the average speed of light is 0.75 times that in a vacuum, and so a body of water, instrumented with light sensors, forms an effective detector for any charged particle travelling faster than this. Because water is cheap and transparent, it lends itself to non-accelerator particle physics experiments where low event rates require huge detectors, particularly for the detection of relativistic electrons from neutrino scattering and proton decay [2,3].

A typical size for this class of detector is a few thousand tonnes, i.e. of order 10 m radius. This requires light detectors which cover a fair fraction of the surface area of the water tank - hundreds of square metres - particularly when looking for low energy events which produce few Cherenkov photons. Two such detectors have been built on a 1000 tonne scale - IMB in Ohio [4] and Kamiokande [5] in Japan. Both were originally designed to search for proton decay, both attained fame by detecting neutrinos from Supernova SN1987A, and now Kamiokande has been modified to lower its energy threshold for the observation of solar neutrinos. Two additional Cherenkov detectors are being built on a scale as large or larger: the Sudbury Neutrino Observatory, with which we are presently concerned, and SuperKamiokande [6], a 50000 tonne light water Cherenkov detector.

At present the only serious option in light detection is an array of photomultiplier tubes (PMTs). It is now possible to obtain PMTs which have relatively large sensitive areas (of the order 1000 cm^2 each) for relatively low cost (several hundred dollars), with good timing (a few ns) and good quantum efficiency (~20%). Moreover, the effective area of a PMT can be significantly improved by using light con-

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[*] Corresponding author. Tel. +44 1865 273380, fax +44 1865 273418, e-mail n.jelley@physics.oxford.ac.uk.

[1] Present address: TANDAR, Departamento de Física, Comisión Nacional de Energía Atómica, Av. del Libertador 8250, 1429 Buenos Aires, Argentina.

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